

Derivation of K_d from a

In what follows we have left off the dependence on wavelength. In addition we have left off effects due to inelastic scattering.

Modeling $K_d(z)$ requires the inclusion of the depth dependence of the shape of the radiance distribution. This can be accomplished by using Gershun's equation:

$$K(z) = a(z) / \bar{\mu}(z)$$

where $a(z)$ is the absorption coefficient and $\bar{\mu}(z)$ is the average cosine of the light field.

$K_d(z)$ differs from $K(z)$ by only a few percent, so that we may set:

$$K_d(z) \approx a(z) / \bar{\mu}(z) \quad .$$

Berwald et al. (1995) have derived a parametric model for the dependence of $\bar{\mu}(z)$ on ω_0 for a vertical sun in a black sky. We will assume the same depth dependence for an ordinary sky. This is not precise, but the average cosine varies slowly and has a typical range of only about 10%. The model is:

$\omega_0 = b / c$, where b and c are the total scattering and attenuation coefficients, including water.

$\tau = cz$, is the optical depth.

$$\bar{\mu}_\infty(\tau) = \bar{\mu}_\infty + (\bar{\mu}(0) - \bar{\mu}_\infty) \exp(-P_\tau \tau).$$

$$\bar{\mu}_\infty = -1.59 \omega_0^4 + 1.71 \omega_0^3 - 0.467 \omega_0^2 - 0.347 \omega_0 + 1$$

$\bar{\mu}(0)$ = cosine of refracted solar zenith angle

$$P_\tau(\omega_0) = -0.166 \omega_0^2 + 0.341 \omega_0 + 0.0305$$

Berwald et. al. Limnology and Oceanography 1995, vol. 40, no8, pp. 1347-1357 .